Analysis, post-processing and visualization tools



Javier Junquera



Andrei Postnikov



Summary of different tools for post-processing and visualization

DENCHAR



PLRHO



DOS, PDOS



MACROAVE

SIESTA's output with XCRYSDEN





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DENCHAR plots the charge density and wave functions in real space

Wave functions

$$\begin{split} \psi_i(\vec{k},\vec{r}\,) &= \sum_{\mu} e^{i\vec{k}\cdot\vec{R}_{\mu}} c_{\mu i}(\vec{k}) \phi_{\mu}(\vec{r}) \\ \uparrow \\ \text{Coefficients of the eigenvector } \psi_i(\vec{k}) & \text{atomic orbitals} \\ \text{with eigenvalue } E_i(\vec{k}) \end{split}$$

Charge density

$$\rho\left(\vec{r}\right) = \sum_{i} \int_{BZ} n_{i}(\vec{k}) |\psi_{i}(\vec{k},\vec{r})|^{2} d\vec{k}$$
$$= \sum_{\mu\nu} \rho_{\mu\nu} \phi_{\nu}^{*}(\vec{r}) \phi_{\mu}(\vec{r})$$

density matrix

DENCHAR operates in two different modes: 2D and 3D

2D

• Charge density and/or electronic wave functions are printed on a regular grid of points contained in a 2D plane specified by the user.

• Used to plot contour maps by means of 2D graphics packages.

3D

 Charge density and/or electronic wave functions are printed on a regular grid of points in 3D.

Results printed in Gaussian Cube format.

• Can be visualized by means of standard programs (Moldel, Molekel, Xcrysden)

How to compile DENCHAR...

\$ cd siesta/Src

make denchar ifort -c -03 f2k

ifort -c -O3 f2kcli.F90 ifort -c -O3 m denchar init.F

ifort -c -O3 m_denchar_geom.f

ifort -c -O3 m_denchar_io.F

ifort -c -O3 m_denchar_neighb.f

ifort -c -O3 m_denchar_work.f

ifort -c -O3 denchar.F

ifort -o denchar \setminus

Use the same arch.make file as for the compilation of serial SIESTA

Versions before 2.0.1, please check for patches in http://fisica.ehu.es/ag/siesta-extra/issues.html

...and where to find the User's Guide and some Examples

\$ cd siesta/Util/Denchar
\$ ls
Docs Examples README-Source

How to run DENCHAR...

SIESTA

WriteDenchar WriteWaveFunctions %block WaveFuncKPoints 0.0 0.0 0.0 %endblock WaveFuncKPoints

.true. .true.

Only if you want to plot wave functions

Output of SIESTA required by DENCHAR SystemLabel.PLD SystemLabel.DIM SystemLabel.DM SystemLabel.WFS (only if wave functions) ChemicalSpecies.ion (one for each chemical species)

DENCHAR

\$ In -s ~/siesta/Src/denchar .
\$ denchar < dencharinput.fdf</pre>

You do not need to rerun SIESTA to run DENCHAR as many times as you want

Input of DENCHAR General issues

- Written in fdf (Flexible Data Format), as in SIESTA
- It shares some input variables with SIESTA
 - SystemLabel NumberOfSpecies ChemicalSpeciesLabel

Some other input variables are specific of DENCHAR (all of them start with "Denchar.")

- To specify the mode of usage
- To define the plane or 3D grid where the charge/wave functions are plotted
- To specify the units of the input/output

Input of DENCHAR can be attached at the end of the input file of SIESTA

Input of DENCHAR How to specify the mode of run

Denchar.TypeOfRun	(string)	2D or 3D
Denchar.PlotCharge	(logical)	.TRUE. or .FALSE.
If .true. SystemLabel.DM m	ust be present	
 Denchar.PlotWaveFunctions 	(logical)	.TRUE. or .FALSE.
If .true. SystemLabel.WFS	must be present	

Either one or the other (or both of them) must be .true.

Input of DENCHAR How to specify the plane

Denchar.PlaneGeneration

NormalVector

TwoLines

ThreePoints

ThreeAtomicIndices

+ more variables to define the

Plane of the plot in 2D mode x-y plane in 3D mode

(string)

generation object (the normal vector, lines, points or atoms) origin of the plane x-axis size of the plane number of points in the grid

• Different variables described in the User Guide (take a look to the Examples)

Output of DENCHAR 2D mode

Charge density

Spin unpolarized:

self-consistent charge (.CON.SCF)

deformation charge (.CON.DEL)

Spin polarized:

density spin up (.CON.UP) density spin down (.CON.DOWN) deformation charge (.CON.DEL) magnetization (.CON.MAG)

Wave functions

Wave function for different bands (each wavefunction in a different file) .CON.WF#, where # is the number of the wf (If spin polarized, suffix .UP or .DOWN)

	Format	
x_i	y_i	$f\left(x_{i},y_{i} ight)$
-5.00000	-5.00000	-0.00560
-5.00000	-4.74359	-0.00402
•	•	•

Output of DENCHAR 3D mode

Charge density

Spin unpolarized:

self-consistent charge (.RHO.cube) deformation charge (.DRHO.cube)

Spin polarized:

density spin up (.RHO.UP.cube) density spin down (.RHO.DOWN.cube) deformation charge (.DRHO.cube)

Wave functions

Wave function for different bands (each wavefunction in a different file) .WF#.cube, where # is the number of the wf (If spin polarized, suffix .UP or .DOWN)

Format

Gaussian Cube format

Atomic coordinates and grid points in the reference frame given in the input Reference frame orthogonal

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PLRHO plots a 3D isosurface of the charge density and colours it with a second function

- $\bullet \, \rho \left(\vec{r} \right)$
- LDOS integrated in a given energy interval
- $ho\left(ec{r}
 ight)$ + electrostatic potential
- $ho\left(ec{r}
 ight)$ + total potential
- $\cdot \rho \left(\vec{r}
 ight)$ + spin density

Plrho reads the values of the functions in the real space grid and Interpolates to plot the 3D surface.

How to compile PLRHO

First you need to install the PGPLOT library, available from

http://www.astro.caltech.edu/~tjp/pgplot

You can find plrho at

~/siesta/Utils/Plrho

Then compile PLRHO with

\$ f90 plrho.f –IX11 –Ipgplot –o plrho

Check plrho_guide.txt for extra information.

How to run PLRHO

SIESTA

SaveRho.true.SaveElectrostaticPotential.true.SaveTotalPotential.true.%block LocalDensityOfStates%block AtomicCoordinatesOrigin

Depending on what you want to plot

If you want to center the system

Output of SIESTA required by PLRHO SystemLabel.RHO SystemLabel.VH SystemLabel.VT SystemLabel.LDOS

PLRHO Prepare the input file plrho.dat \$ plrho

You do not need to rerun SIESTA to run PLRHO as many times as you want

'h2o'	System label, used to name input files
'vh'	<pre>Function(s) to plot ('rho' 'ldos' 'spin' 'vt' 'vh')</pre>
90.0 0.0 -90.0	Euler rotation angles alpha, beta, gamma
4.e-3	Value of electron density (or LDOS) for isosurface
-0.08 +0.01 +0.12	Saturation range for color function
'unformatted'	Data format ('formatted' 'unformatted')
'/CPS' (Dutput option ('/XWIN' '/PS' '/CPS' +more)

'h2o'	System label, used to name input files
'vh'	<pre>Function(s) to plot ('rho' 'ldos' 'spin' 'vt' 'vh')</pre>
<u>9</u> 0.0 0.0 -90.0	Euler rotation angles alpha, beta, gamma
4.e-3	Value of electron density (or LDOS) for isosurface
-0.08 +0.01 +0.12	Saturation range for color function
'unformatted'	Data format ('formatted' 'unformatted')
'/CPS'	Output option ('/XWIN' '/PS' '/CPS' +more)

'rho'	$ ho\left(ar{r} ight)$
'ldos'	LDOS integrated in a given energy interval
'vh'	$ ho\left(ec{r} ight)$ + electrostatic potential
'vt'	$ ho\left(ec{r} ight)$ + total potential
'spin'	$ ho\left(ec{r} ight)$ + spin density

'h2o'	System label, used to name input files
'vh'	Function(s) to plot ('rho' 'ldos' 'spin' 'vt' 'vh')
90.0 0.0 -90.0	Euler rotation angles alpha, beta, gamma
4.e-3	Value of electron density (or LDOS) for isosurface
-0.08 +0.01 +0.	12 Saturation range for color function
'unformatted'	Data format ('formatted' 'unformatted')
'/CPS'	Output option ('/XWIN' '/PS' '/CPS' +more)

Viewpoint is always from above (positive z axis)



To view the system from a different angle, rotate it with the Euler angles

'h2o'		System label, used to name input files
'vh'		Function(s) to plot ('rho' 'ldos' 'spin' 'vt' 'vh')
90.0	0.0 -90.0	Euler rotation angles alpha, beta, gamma
4.e-3		Value of electron density (or LDOS) for isosurface
-0.08	+0.01 +0.12	Saturation range for color function
'unfor	matted'	Data format ('formatted' 'unformatted')
'/CPS'		Dutput option ('/XWIN' '/PS' '/CPS' +more)

Example: view from -y (Euler angles = 90 -90 -90)



'h2o'	System label, used to name input files
'vh'	Function(s) to plot ('rho' 'ldos' 'spin' 'vt' 'vh')
90.0 0.0 -90.0	Euler rotation angles alpha, beta, gamma
4.e-3	Value of electron density (or LDOS) for isosurface
-0.08 +0.01 +0.1	2 Saturation range for color function
'unformatted'	Data format ('formatted' 'unformatted')
'/CPS'	Output option ('/XWIN' '/PS' '/CPS' +more)

Example: view from –y (Euler angles = 90 -90 -90)

Reference axes System axes



'h2o'	System label, used to name input files
'vh'	<pre>Function(s) to plot ('rho' 'ldos' 'spin' 'vt' 'vh')</pre>
90.0 0.0 -90.0	Euler rotation angles alpha, beta, gamma
4.e-3	Value of electron density (or LDOS) for isosurface
-0.08 +0.01 +0.12	Saturation range for color function
'unformatted'	Data format ('formatted' 'unformatted')
'/CPS'	Output option ('/XWIN' '/PS' '/CPS' +more)

Example: view from –y (Euler angles = 90 -90 -90)

Reference axes System axes



Alpha: first rotation around z

'h2o'	System label, used to name input files
'vh'	<pre>Function(s) to plot ('rho' 'ldos' 'spin' 'vt' 'vh')</pre>
90.0 0.0 -90.0	Euler rotation angles alpha, beta, gamma
4.e-3	Value of electron density (or LDOS) for isosurface
-0.08 +0.01 +0.12	Saturation range for color function
'unformatted'	Data format ('formatted' 'unformatted')
'/CPS' (Output option ('/XWIN' '/PS' '/CPS' +more)

Example: view from –y (Euler angles = 90 -90 -90)

Reference axes System axes



'h2o'	System label, used to name input files
'vh'	<pre>Function(s) to plot ('rho' 'ldos' 'spin' 'vt' 'vh')</pre>
90.0 0.0 -90.0	Euler rotation angles alpha, beta, gamma
4.e-3	Value of electron density (or LDOS) for isosurface
-0.08 +0.01 +0.12	Saturation range for color function
'unformatted'	Data format ('formatted' 'unformatted')
'/CPS'	Dutput option ('/XWIN' '/PS' '/CPS' +more)

Example: view from -y (Euler angles = 90 -90 -90)

Reference axes System axes



Gamma: second rotation around z

'h2o'	System label, used to name input files
'vh'	<pre>Function(s) to plot ('rho' 'ldos' 'spin' 'vt' 'vh')</pre>
90.0 0.0 -90.0	Euler rotation angles alpha, beta, gamma
4.e-3	Value of electron density (or LDOS) for isosurface
-0.08 +0.01 +0.12	Saturation range for color function
'unformatted'	Data format ('formatted' 'unformatted')
'/CPS' (Output option ('/XWIN' '/PS' '/CPS' +more)



Output of PLRHO



Output of PLRHO

'h2o'	System label, used to name input files
'vh'	<pre>Function(s) to plot ('rho' 'ldos' 'spin' 'vt' 'vh')</pre>
90.0 0.0 -90.0	Euler rotation angles alpha, beta, gamma
4.e-3	Value of electron density (or LDOS) for isosurface
-0.08 +0.01 +0.12	Saturation range for color function
'unformatted'	Data format ('formatted' 'unformatted')
'/CPS'	Output option ('/XWIN' '/PS' '/CPS' +more)



H₂O molecule

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Density Of States (DOS) g(E)dE the number of one-electron levels between *E* and *E* + *dE*

Si bulk



Density Of States (DOS) g(E)dE the number of one-electron levels between *E* and *E* + *dE*

0 n E_{F} -5 -5 E (eV) E (eV) -10 -10 -15 -15 -20∟ K -20 Γ Х W Г DOS (eV^{-1})

$$g(E) = rac{1}{N_{ec k}}\sum_{i}^{bands}\sum_{ec k}\delta\left(E-E_i(ec k)
ight)$$

Units: (Energy)⁻¹

Si bulk

Projected Density Of States (PDOS) $g_{\mu}(E)dE$ the number of one-electron levels with weight on orbital μ between E and E + dE

$$g_{\mu}(E) = \frac{1}{N_{\vec{k}}} \sum_{i} \sum_{i} \sum_{\nu} c_{\nu i}^{*}(\vec{k}) c_{\mu i}(\vec{k}) S_{\nu \mu}(\vec{k}) \delta\left(E - E_{i}(\vec{k})\right)$$
Coefficients of the eigenvector $\psi_{i}(\vec{k})$ Overlap matrix of the atomic basis with eigenvalue $E_{i}(\vec{k})$

Units: (Energy)⁻¹

Relation between the DOS and PDOS:

$$g(E) = \sum_{\mu} g_{\mu}(E)$$

The eigenvalues are broadening by a gaussian to smooth the shape of the DOS and PDOS

 $\delta\left(E - E_i(\vec{k})\right) \to \frac{1}{\sigma\sqrt{\pi}} e^{-\frac{\left(E - E_i(\vec{k})\right)^2}{\sigma^2}}$



Two step procedure to produce smooth DOS and PDOS

First:

Run a simulation with a sensible (converged) number of k-points kgrid_cutoff %block kgrid_Monkhorst_Pack Get converged geometry and density matrix Second:

Starting from the previously converged geometry and density matrix, run a single SCF step with fixed geometry, with many more k-points

DM.UseSaveDM	.true.
MaxSCFIterations	1
MD.NumCGsteps (or equivalent)	0
Increase number of k-points (see above)	
%block ProjectedDensityOfStates	

%block ProjectedDensityOfStates -20.0 10.0 0.200 500 eV %endblock ProjectedDensityOfStates

-20.0 10.0 : Energy window where the DOS and PDOS will be computed

%block ProjectedDensityOfStates -20.0 10.0 0.200 500 eV %endblock ProjectedDensityOfStates

-20.0 10.0 : Energy window where the DOS and PDOS will be computed

0.200 : Peak width of the gaussian used to broad the eigenvalues (energy)

%block ProjectedDensityOfStates -20.0 10.0 0.200 500 eV %endblock ProjectedDensityOfStates

-20.0 10.0 : Energy window where the DOS and PDOS will be computed

0.200 : Peak width of the gaussian used to broad the eigenvalues (energy)

500 : Number of points in the histogram

%block ProjectedDensityOfStates -20.0 10.0 0.200 500 eV %endblock ProjectedDensityOfStates

-20.0 10.0 : Energy window where the DOS and PDOS will be computed

0.200 : Peak width of the gaussian used to broad the eigenvalues (energy)

500 : Number of points in the histogram

eV : Units in which the previous energies are introduced

Output for the Density Of States

SystemLabel.DOS

Format

Energy (eV)	DOS Spin Up (eV ⁻¹)	DOS Spin Down (eV ⁻¹)		
-15.99598	0.03075	0.00000		
-15.97596	0.03580	0.00000		
-15.95594	0.04115	0.00000		

Output for the Projected Density Of States

SystemLabel.PDOS

```
<pdos>
<nspin>1</nspin>
<norbitals> 26</norbitals>
<energy_values units="eV">
           -19.99998
           -19.97996
           -19.95994
</energy_values>
<orbital
index="1"
atom index="1"
species="Si"
position=" 0.000000 0.000000
                                    0.000000"
n="3"
 1=" 0"
 m=" 0"
z="1"
>
<data>
  0.00000
  0.00000
  0.00000
</data>
</orbital>
</pdos>
```

Written in XML

Energy Window

One element <orbital> for every atomic orbital in the basis set

How to digest the SystemLabel.PDOS file

During the compilation of SIESTA

For some compilers, the libwxml.a library needs to be compiled with "-DWXML_INIT_FIX" (see known issues in http://fisica.ehu.es/ag/siesta-extra/issues.html)

pdosxml (by Alberto García)

Siesta/Util/pdosxml

Edit the readme file to:

Learn how to select the orbitals whose PDOS will be accumulated

How to compile the code

How to run the code

fmpdos (by Andrei Postnikov)

Download from

http://www.home.uni-osnabrueck.de/apostnik/download.html

Compile and follow the instructions at run-time

Normalization of the DOS and PDOS

$$\int_{-\infty}^{+\infty} g(E) \, dE =$$

Number of bands ____ Number of atomic orbitals per k-point in the unit cell

 $\int_{-\infty}^{+\infty} g(E) n(E) \, dE = \begin{array}{c} \text{Number of electrons} \\ \text{in the unit cell} \end{array}$

Occupation factor at energy *E*

Example of DOS and PDOS

$$g_{\mu}(E) = \frac{1}{N_{\vec{k}}} \sum_{i}^{bands} \sum_{\vec{k}} \sum_{\nu} c_{\nu i}^{*}(\vec{k}) c_{\mu i}(\vec{k}) S_{\nu \mu}(\vec{k}) \delta\left(E - E_{i}(\vec{k})\right)$$



J. Junquera *et al.* Surf. Sci. 482-485, 625 (2001) SIESTA, single-zeta polarized basis



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How to extract from the immense detail provided by firstprinciples calculations on surfaces reliable values of the physical quantities of interest



Physical quantities of interest in surfaces and interfaces

- Charge densities at the surface/interface
- Dipole moment densities at the surface/interface
- Work functions (surfaces) and band offsets (interfaces)



 $VBO = \Delta E_{oldsymbol{v}} + \Delta V$

 ΔE_v : the band structure term

Difference of the top of the valence bands

From two independent bulk band structure calculations of the bulk material

ΔV : jump of the average electrostatic potential

Contains all the intrinsic interface effects

Obtained by nanosmoothing the electrostatic potential

First step: average in the plane

 $\overline{V}(z) = rac{1}{S} \int_{S} dx dy \,\, V\left(ec{r}
ight)$



Second step: nanosmooth the planar $\overline{V}(z) = \int dz' f(z-z')\overline{V}(z')$ average on the *z*-direction **BaTiO**₃ vacuum vacuum 0.0 -5.0 electrostatic potential (eV) -10.0 -15.0 -20.0 -25.020 ~ 0 ~ 0 ~ 0

Atomic scale fluctuations are washed out by filtering the magnitudes via convolution with smooth functions

$$ar{
ho}(z) = \int dz' f(z-z') ar{
ho}(z')
onumber \ f(z-z') = \int dz'' \omega_{l_1}(z-z'') \omega_{l_2}(z''-z')$$



ΔV is readily obtained from the nanosmoothed potential



How to compile MACROAVE...

```
$ cd ~/siesta/Util/Macroave/Src/
$ make
ifort -c -O3 defs_basis.f90
ifort -c -O3 defs_common.f90
ifort -c -O3 io.f
ifort -c -O3 iorho.f
ifort -c -O3 paste.f
ifort -c -O3 macroave.f
ifort -c -O3 thetaft.f
ifort -c -O3 surpla.f
ifort -c -O3 volcel.f
ifort -c -O3 recipes.f
ifort -c -O3 hdr io.f90
ifort -o macroave.x \
        defs_basis.o defs_common.o io.o iorho.o paste.o macroave.o thetaft.o
        surpla.o volcel.o recipes.o hdr_io.o
```

Use the same arch.make file as for the compilation of serial SIESTA

...and where to find the User's Guide and some Examples

\$ cd ~/siesta/Util/Macroave
\$ ls
Docs Examples Src

How to run MACROAVE

SIESTA

SaveRho	
SaveTotalCharge	
SavelonicCharge	
SaveDeltaRho	
SaveElectrostaticPotential	
SaveTotalPotential	

.true. .true. .true. .true. .true. .true.

Depending on what you want to nanosmooth

Output of SIESTA required by MACROAVE SystemLabel.RHO SystemLabel.TOCH SystemLabel.IOCH SystemLabel.DRHO SystemLabel.VH SystemLabel.VT

MACROAVE

Prepare the input file macroave.in \$ ~/siesta/Util/Macroave/Src/macroave.x You do not need to rerun SIESTA to run MACROAVE as many times as you want

Siesta	#	Which code have you used to get the input data?
Potential	#	Quantity that will be nanosmoothed
Interface	#	Name of the file where the input data is stored
2	#	Number of convolutions required to calculate the macro. ave.
7.47622255	#	First length for the filter function in macroscopic average
7.56114385	#	Second length for the filter function in macroscopic average
500	#	Total number of electrons
spline	#	Type of interpolation

The same code with the same input runs with information provided by

Siesta

or

Abinit

(indeed it should be quite straight forward to generalize to any other code)

Siesta	#	Which code have you used to get the input data?
Potential	#	Quantity that will be nanosmoothed
Interface	#	Name of the file where the input data is stored
2	#	Number of convolutions required to calculate the macro. ave.
7.47622255	#	First length for the filter function in macroscopic average
7.56114385	#	Second length for the filter function in macroscopic average
500	#	Total number of electrons
spline	#	Type of interpolation

Name of the magnitude that will be nanosmoothed

Potential:	SystemLabel.VH
Charge:	SystemLabel.RHO
TotalCharge:	SystemLabel.TOCH

Siesta	#	Which code have you used to get the input data?
Potential	#	Quantity that will be nanosmoothed
Interface		Name of the file where the input data is stored
2	#	Number of convolutions required to calculate the macro. ave.
7.47622255	#	First length for the filter function in macroscopic average
7.56114385	#	Second length for the filter function in macroscopic average
500	#	Total number of electrons
spline	#	Type of interpolation

SystemLabel

Siesta	#	Which code have you used to get the input data?
Potential	#	Quantity that will be nanosmoothed
Interface	#	Name of the file where the input data is stored
2	alide alide	Number of convolutions required to calculate the macro. ave.
7.47622255	#	First length for the filter function in macroscopic average
7.56114385	#	Second length for the filter function in macroscopic average
500	#	Total number of electrons
spline	#	Type of interpolation

Number of square filter functions used for nanosmoothing

1	Surfaces
2	Interfaces and superlattices

Siesta	#	Which code have you used to get the input data?
Potential	#	Quantity that will be nanosmoothed
Interface	#	Name of the file where the input data is stored
2	#	Number of convolutions required to calculate the macro. ave.
7.47622255	#	First length for the filter function in macroscopic average
7.56114385	井	Second length for the filter function in macroscopic average
500	#	Total number of electrons
spline	#	Type of interpolation

Length of the different square filter functions (in bohrs)



Siesta	#	Which code have you used to get the input data?
Potential	#	Quantity that will be nanosmoothed
Interface	#	Name of the file where the input data is stored
2	#	Number of convolutions required to calculate the macro. ave.
7.47622255	#	First length for the filter function in macroscopic average
7.56114385	#	Second length for the filter function in macroscopic average
500	#	Total number of electrons
spline	#	Type of interpolation

Total number of electrons

(used only to renormalize if we nanosmooth the electronic charge)

Siesta	# Which code have you used to get the input data?
Potential	# Quantity that will be nanosmoothed
Interface	# Name of the file where the input data is stored
2	# Number of convolutions required to calculate the macro. ave.
7.47622255	# First length for the filter function in macroscopic average
7.56114385	# Second length for the filter function in macroscopic average
500	# Total number of electrons
spline	# Type of interpolation
2 7.47622255 7.56114385 500 spline	# Number of convolutions required to calculate the macro. ave. # First length for the filter function in macroscopic average # Second length for the filter function in macroscopic average # Total number of electrons # Type of interpolation

Type of interpolation from the SIESTA mesh to a FFT mesh

Spline or Linear

Output of MACROAVE

Planar average

$$\overline{
ho}(z) = rac{1}{S} \int_{S} dx dy \
ho \left(ec{r}
ight)$$

SystemLabel.PAV

Nanosmoothed

$$\overline{\overline{
ho}}(z) = \int dz^{'} f(z-z^{'}) \overline{
ho}(z^{'})$$

SystemLabel.MAV

|--|

z		average
0.00000000000		0.000000000000
0.000095367432		0.000000000000
0.000190734863		0.000000000000
	•	

Units

Coordinates: bohr

Potential: eV

Charge density: electrons/bohr³

To learn more on nanosmoothing and how to compute work functions and band offsets with SIESTA

IOP PUBLISHING

JOURNAL OF PHYSICS: CONDENSED MATTER

J. Phys.: Condens. Matter 19 (2007) 213203 (34pp)

doi:10.1088/0953-8984/19/21/213203

TOPICAL REVIEW

Nanoscale smoothing and the analysis of interfacial charge and dipolar densities

Javier Junquera^{1,2}, Morrel H Cohen^{2,3} and Karin M Rabe²

PHYSICAL REVIEW B 67, 155327 (2003)

First-principles calculation of the band offset at BaO/BaTiO₃ and SrO/SrTiO₃ interfaces

Javier Junquera,¹ Magali Zimmer,¹ Pablo Ordejón,² and Philippe Ghosez¹